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ON THE VALIDATION OF A SPECTRAL/SPATIAL CBIR SYSTEM FOR HYPERSPECTRAL IMAGES

Miguel A. Veganzones, Manuel Graña

Grupo Inteligencia Computacional Universidad del País Vasco

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CBIR systems

- Images indexed by feature vectors extracted by means of computer vision and digital image processing techniques.
 - The interrogation to the database is done through the presentation of a query image.
 - The answer are the most similar images in the database according to some similarity measure.

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Remote Sensing CBIR systems

- Motivation: huge amount of Earth Observation data provided by remote sensors.
- Approaches to CBIR in remote sensing images proposed up to now are focused on panchromatic, SAR or low dimension multispectral images.
- CBIR techniques has not been properly addressed for the case of hyperspectral images.

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Hyperspectral CBIR systems

- Few works in the literature.
- Image features as the endmembers induced by some Endmembers Induction Algorithm (EIA).
- Inconvenient: they can not discriminate among images with the same induced endmembers but very different spatial distributions.
- We propose an Spectral-Spatial CBIR system for hyperspectral databases.

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CBIR systems validation

- Most frequently used evaluation measures:
 - Precision (p): fraction of the returned images that are relevant to the query.
 - Recall (q): fraction of returned relevant images respect to the total number of relevant images in the database.
- Implementation:
 - A-priori ground-truth knowledge (categories).
 - User's online evaluation.

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Validation in Remote Sensing domain

- Handicaps:
 - Lack of ground-truth knowledge (categories).
 - Users have difficulties to evaluate the system's response.
- We validate the proposed Spectral-Spatial CBIR system using synthetic hyperspectral data.
- We are working on CBIR systems validation without a-priori ground-truth knowledge.

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Linear Mixing Model

LMM

- $H = A \cdot E + \eta$
- $\mathbf{h}(\mathbf{x}, \mathbf{y}) = a_1(x, y) \cdot \mathbf{e_1} + a_2(x, y) \cdot \mathbf{e_2} + \ldots + a_p(x, y) \cdot \mathbf{e_p} + \eta$

where:

- *H* is an hyperspectral image.
- E is the set of materials spectral signatures (endmembers): spectral information.
- A is the set of fractional abundance images: spatial information.
- η is additive noise.

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Spectral/Spatial features

- We characterize an hyperspectral image H_{α} by a tuple $\langle E_{\alpha}, \Phi_{\alpha} \rangle$, where:
 - $E_{\alpha} = \{\mathbf{e}_{1}^{\alpha}, \dots, \mathbf{e}_{p_{\alpha}}^{\alpha}\}$ is the set of p_{α} induced endmembers. • $\Phi_{\alpha} = \{\phi_{1}^{\alpha}, \dots, \phi_{p_{\alpha}}^{\alpha}\}$ is the set of fractional abundance maps.
- An EIA is used to induce the spectral signatures (the endmembers) of the image.
- An unmixing method extracts from the image the fractional abundances of each endmember.

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Spectral/Spatial CBIR system

- Image similarity: similarity between image spectra and their relative abundance proportions.
- Spectral-Spatial dissimilarity: is a version of the Integrated Region Matching (IRM) dissimilarity function used for region matching-based image retrieval.
- Depends on two aspects:
 - The similarity between each region of the two images (spectral similarity).
 - The significance of each region matching (spatial similarity).

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Spectral-Spatial dissimilarity function

• Given two images, H_{α} and H_{β} :

We compute the Spectral Distance Matrix, D_{α.β}:

$$D_{lpha,eta} = [d_{ij}; i = 1, \dots, p_{lpha}; j = 1, \dots, p_{eta}]$$

where d_{ij} is a distance between the endmembers \mathbf{e}_i^{α} and \mathbf{e}_i^{β} . Then, the Spectral-Spatial dissimilarity function, $s(H_{\alpha}, H_{\beta})$, is defined as:

$$s\left(H_{\alpha}, H_{\beta}\right) = \sum_{i,j} r_{ij} d_{ij}$$

where r_{ij} is the significance associated to $d_{ij}.$ 101000100010101010107

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Significance matrix

- Thus, the problem reduces to choosing the significance matrix $R_{\alpha,\beta} = [r_{ij}; i = 1, \dots, p_{\alpha}; j = 1, \dots, p_{\beta}].$
- We followed the most similar highest priority (MSHP) principle, making use of the average abundances $\bar{\Phi}_{\alpha}$ and $\bar{\Phi}_{\beta}$.
- The average abundances represent "significance credits" assigned to the spectral distances by IRM algorithm.

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IRM algorithm

• Set $\mathcal{L} = \{\}$ and denote $\mathcal{M} = \{(i,j): i = 1, \ldots, p_{\alpha}; j = 1, \ldots, p_{\beta}\}.$ 2 Choose the minimum d_{ij} for $(i,j) \in \mathcal{M} - \mathcal{L}$. Label the corresponding (i, j) as (i', j'). $r_{i'j'} = \min\left(\bar{\phi}_{i'}^{\alpha}, \bar{\phi}_{j'}^{\beta}\right)$ If $\bar{\phi}_{i'}^{\alpha} < \bar{\phi}_{i'}^{\beta}$, set $r_{i'j} = 0$, $j \neq j'$; otherwise, set $r_{ij'} = 0$, $i \neq i'$. If $\bar{\phi}_{i'}^{\alpha} < \bar{\phi}_{i'}^{\beta}$, set $\bar{\phi}_{i'}^{\alpha} = 0$ and $\bar{\phi}_{i'}^{\beta} = \bar{\phi}_{i'}^{\beta} - \bar{\phi}_{i'}^{\alpha}$; otherwise, set $ar{\phi}^{eta}_{i'}=0$ and $ar{\phi}^{lpha}_{i'}=ar{\phi}^{lpha}_{i'}-ar{\phi}^{eta}_{j'}.$ • If $\sum_{i=1}^{p_{\alpha}} \bar{\phi}_i^{\alpha} > 0$ and $\sum_{i=1}^{p_{\beta}} \bar{\phi}_i^{\beta} > 0$, go to step 2; otherwise, stop.

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Synthesis procedure

- Generated as linear mixtures of a set of spectra (ground-truth endmembers) and synthesized abundance coefficients (ground-truth abundances).
 - Ground-truth endmembers: randomly selected from a subset of the USGS spectral library.
 - Ground-truth abundances: Gaussian Random Fields with Matern correlation function of parameters $\theta_1 = 10$ and $\theta_2 = 1$.
- We ensure that there are regions of almost pure endmembers and that the abundance coefficients sum up to one.

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Synthesis example



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Synthetic datasets

- 18000 hyperspectral images divided in nine datasets of 2000 images each.
- Each dataset is characterized by:
 - One of three pools of groundtruth endmembers, with 5, 10 and 20 endmembers each, representing an increasing diversity in the materials.
 - One of three spatial sizes, with images having $64\times 64,$ 128×128 and 256×256 pixels, representing different spatial scales.
- All the synthesized hyperspectral images have 269 spectral bands *per* pixel.
- Each image is built with 2 to 5 endmembers randomly selected from the corresponding pool of available groundtruth endmembers.

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Experiments

- We have performed independent experiments over each of the nine synthetic hyperspectral datasets using the proposed Spectral-Spatial dissimilarity function where:
 - The distance between endmembers is measured by the Euclidean distance, s_{euc}, and the Spectral Angle Map pseudo-distance, s_{sam}.
 - For each image, we apply independently the N-FINDER and the EIHA endmember induction algorithms to induce the set of endmembers.
- The abundances have been calculated by Full-Constrained Least Squares Unmixing (FCLSU).
- The Matlab code for the hyperspectral image synthesis and endmember induction is available from http://www.ehu.es/ccwintco/index.php/GIC-source-code-freelibre.

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Experiment definition

• For each image H_{α} in a dataset we calculate the dissimilarity between H_{α} and each of the images in the dataset (N = 2000):

$$\mathbf{s}_{\alpha} = [s_{\alpha 1}, \dots, s_{\alpha N}]$$

where $s_{\alpha,\beta}$ is the dissimilarity between the images H_{α} and H_{β} . • Let us distinguish between:

- $\mathbf{s}_{\alpha}^{\mathrm{GT}}$: the vector of dissimilarities computed using the known ground-truth endmembers.
- s^{IND}_α: the vector of dissimilarities computed using the endmembers induced by one of the EIAs (either N-FINDER or EIHA).

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Experiments ranking definition

 We sort the components of s_α in increasing order, and the resulting shuffled image indices constitute the ranking:

$$\Omega_{\alpha} = [\omega_{\alpha,p} \in \{1,\ldots,N\}; p = 1,\ldots,N]$$

so that $s_{\alpha,\omega_{\alpha,p}} \leq s_{\alpha,\omega_{\alpha,p+1}}$.

• We distinguish rankings $\Omega^{\rm GT}_{lpha}$ and $\Omega^{\rm IND}_{lpha}$ corresponding to the ground-truth and induced dissimilarities, respectively.

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Experiments query definition

- A query Q_k(H_α) is formulated as a search for the k most similar (less dissimilar) images H_β in the dataset with respect to the image H_α, with 1 ≤ k ≤ N.
- The set of returned images $T_k(H_\alpha)$ and the set of relevant images $V_k(H_\alpha)$ for a query $Q_k(H_\alpha)$ are defined as follows:

$$T_{k}(H_{\alpha}) = \Omega_{\alpha,k}^{\text{IND}} = \left[\omega_{\alpha,p}^{\text{IND}} \text{ s.t. } s_{\alpha,\omega_{\alpha,p}}^{\text{IND}} \le s_{\alpha,\omega_{\alpha,k}}^{\text{IND}} \right]$$
(1)
$$V_{k}(H_{\alpha}) = \Omega_{\alpha,k}^{\text{GT}} = \left[\omega_{\alpha,p}^{\text{GT}} \text{ s.t. } s_{\alpha,\omega_{\alpha,p}}^{\text{GT}} \le t \right]$$
(2)

where $t=\bar{s}_{\alpha}^{\rm GT}-2\sigma_{{\rm s}_{\alpha}^{\rm GT}}$, and $\bar{s}_{\alpha}^{\rm GT}$ and $\sigma_{{\rm s}_{\alpha}^{\rm GT}}$ are respectively the mean and standard deviation of ${\rm s}_{\alpha}^{\rm GT}$.

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Experiments quality meassures

• The Precision $P_k(H_{\alpha})$ and Recall $R_k(H_{\alpha})$ for a query $Q_k(H_{\alpha})$ are defined as:

$$P_k(H_\alpha) = \frac{|V_k(H_\alpha) \cap T_k(H_\alpha)|}{|T_k(H_\alpha)|} \text{ and } R_k(H_\alpha) = \frac{|V_k(H_\alpha) \cap T_k(H_\alpha)|}{|V_k(H_\alpha)|}$$

• The Average Precision and Recall of the system for a query of size k are defined as:

$$P_k = \frac{1}{N} \sum_{\alpha=1}^{N} P_k(H_\alpha) \text{ and } R_k = \frac{1}{N} \sum_{\alpha=1}^{N} R_k(H_\alpha)$$

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Experiments quality meassures II

 As summary performance quantity, we calculate the normalized average rank of relevant images:

$$Rank\left(H_{\alpha}\right) = \frac{1}{NN_{\alpha}} \left(\sum_{i=1}^{N_{\alpha}} R_{i} - \frac{N_{\alpha}\left(N_{\alpha} - 1\right)}{2}\right)$$

where N_{α} is the number of relevant images for the query, and R_i is the rank at which the *i*-th image is retrieved.

- This measure is 0 for perfect performance, and approaches 1 as performance worses.
- The average normalized rank ANR for the full dataset is given by:

$$ANR = \frac{1}{N} \sum_{\alpha=1}^{N} Rank \left(H_{\alpha} \right)$$

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Precision-recall curves



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ANR results

			ANR		
Spatial size	Distance	EIA	5-dataset	10-dataset	20-dataset
64×64	Euclidean	EIHA	0.0383	0.0442	0.0356
64×64	Euclidean	N-FINDER	0.0051	0.0120	0.0109
64×64	SAM	EIHA	0.0512	0.0559	0.0558
64×64	SAM	N-FINDER	0.0035	0.0145	0.0318
128×128	Euclidean	EIHA	0.0216	0.0306	0.0228
128×128	Euclidean	N-FINDER	0.0056	0.0108	0.0118
128×128	SAM	EIHA	0.0371	0.0440	0.0458
128×128	SAM	N-FINDER	0.0026	0.0153	0.0340
256×256	Euclidean	EIHA	0.0116	0.0186	0.0189
256×256	Euclidean	N-FINDER	0.0035	0.0119	0.0119
256×256	SAM	EIHA	0.0220	0.0368	0.0412
256×256	SAM	N-FINDER	0.0019	0.0180	0.0316

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Summary

- The proposed spectral-spatial CBIR system is robust to the choice of EIA and between-endmembers distance function.
- We find that the N-FINDER has obtained better results in the sense that its precision-recall curves is systematically above the ones corresponding to the EIHA.

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Future work

- - Develop a retrieval feedback process:
 - Hard due to users limitations to evaluate hyperspectral data by visual inspection.
 - Try new methodologies for CBIR systems validation without a-priori knowledge.
 - Presented in ESA-EUSC-JCR 2011 conference in lspra (ltaly).

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Thanks for your attention

- Contact:
 - Miguel Angel Veganzones
 - Grupo de Inteligencia Computacional
 - Universidad del País Vasco (Spain)
 - E-mail: miguelangel.veganzones@ehu.es
 - Web page: http://www.ehu.es/computationalintelligence

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